

Messier 87

&

Most detailed image of a Black Hole ever obtained with

Event Horizon Telescope (EHT)

Some useful links:

EU Research press conference (Brussels): <https://www.youtube.com/watch?v=Dr20f19czeE>

Press conference at National Science Foundation: <https://www.youtube.com/watch?v=lnJi0Jy692w>

Detailed explanation of the observed phenomenon: https://www.youtube.com/watch?v=vlgs2_C-EtM

Short description of Event Horizon Telescope: <https://www.youtube.com/watch?v=gK-R6U5KCjE>

Live press conference in Italy: <https://www.youtube.com/watch?v=uEhpjDjifGM>

ESO science release: <https://www.eso.org/public/news/eso1907/>

The cluster of galaxies closest to the Milky Way

The Virgo Cluster

Distance: $d = 20$ Mpc

Binding mass: $M \sim 1.2 \times 10^{15} M_{\odot}$

Size: ~ 3 Mpc

About 1500 galaxies (mix of elliptical, S0 & spiral galaxies)



Radio galaxy M87 with jet

Apparent magnitude: $m_V = 9.59$

Distance: 16.40 ± 0.50 Mpc

Total mass: $M_{\text{tot}} = 6 \times 10^{12} M_{\odot}$

Mass central black hole: $M_{\text{BH}} = 7 \times 10^9 M_{\odot}$

Messier 87: one of the most massive galaxies in the *Local Universe*



Radio galaxy *M87* with jet

Apparent magnitude: $m_V = 9.59$

Distance: 16.40 ± 0.50 Mpc

Total mass: $M_{\text{tot}} = 6 \times 10^{12} M_{\odot}$

Mass central black hole: $M_{\text{BH}} = 7 \times 10^9 M_{\odot}$

Messier 87: one of the most massive galaxies in the *Local Universe*



Radio galaxy *M87* with jet

Apparent magnitude: $m_V = 9.59$

Distance: 16.40 ± 0.50 Mpc

Total mass: $M_{\text{tot}} = 6 \times 10^{12} M_{\odot}$

Mass central black hole: $M_{\text{BH}} = 7 \times 10^9 M_{\odot}$

Messier 87: radio galaxy with jet

Distance: 16.4 ± 0.5 Mpc

Total mass: $M_{tot} = 6 \times 10^{12} M_{\odot}$

Mass central black hole: $M_{BH} = 7 \times 10^9 M_{\odot}$

(one of the highest mass known for a black hole)

Optical image



Messier 87: radio galaxy with jet

Distance: 16.4 ± 0.5 Mpc

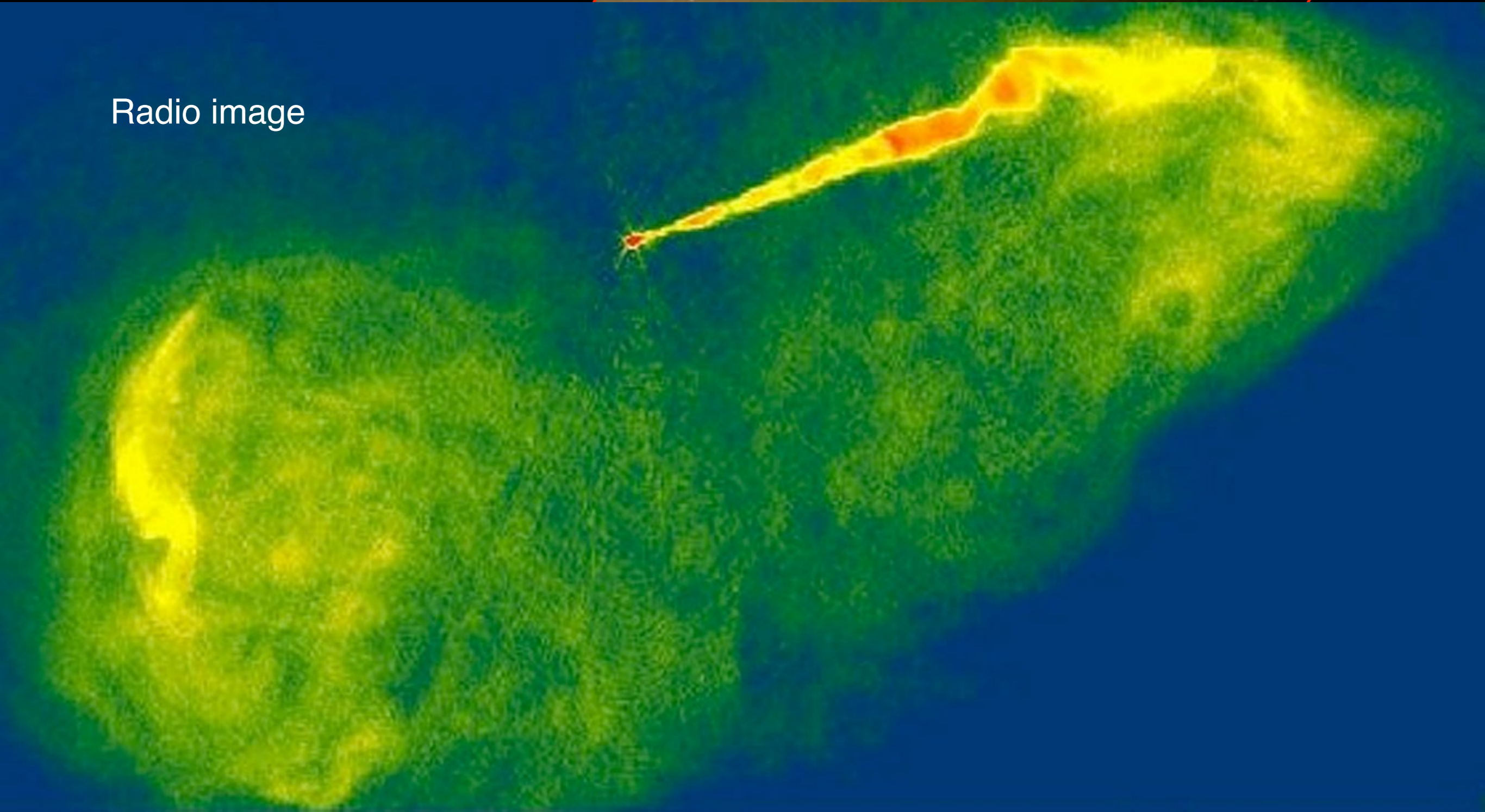
Total mass: $M_{\text{tot}} = 6 \times 10^{12} M_{\odot}$

Mass central black hole: $M_{\text{BH}} = 7 \times 10^9 M_{\odot}$

(one of the highest mass known for a black hole)

Optical image

Radio image



Messier 87: radio galaxy with jet

Distance: 16.4 ± 0.5 Mpc

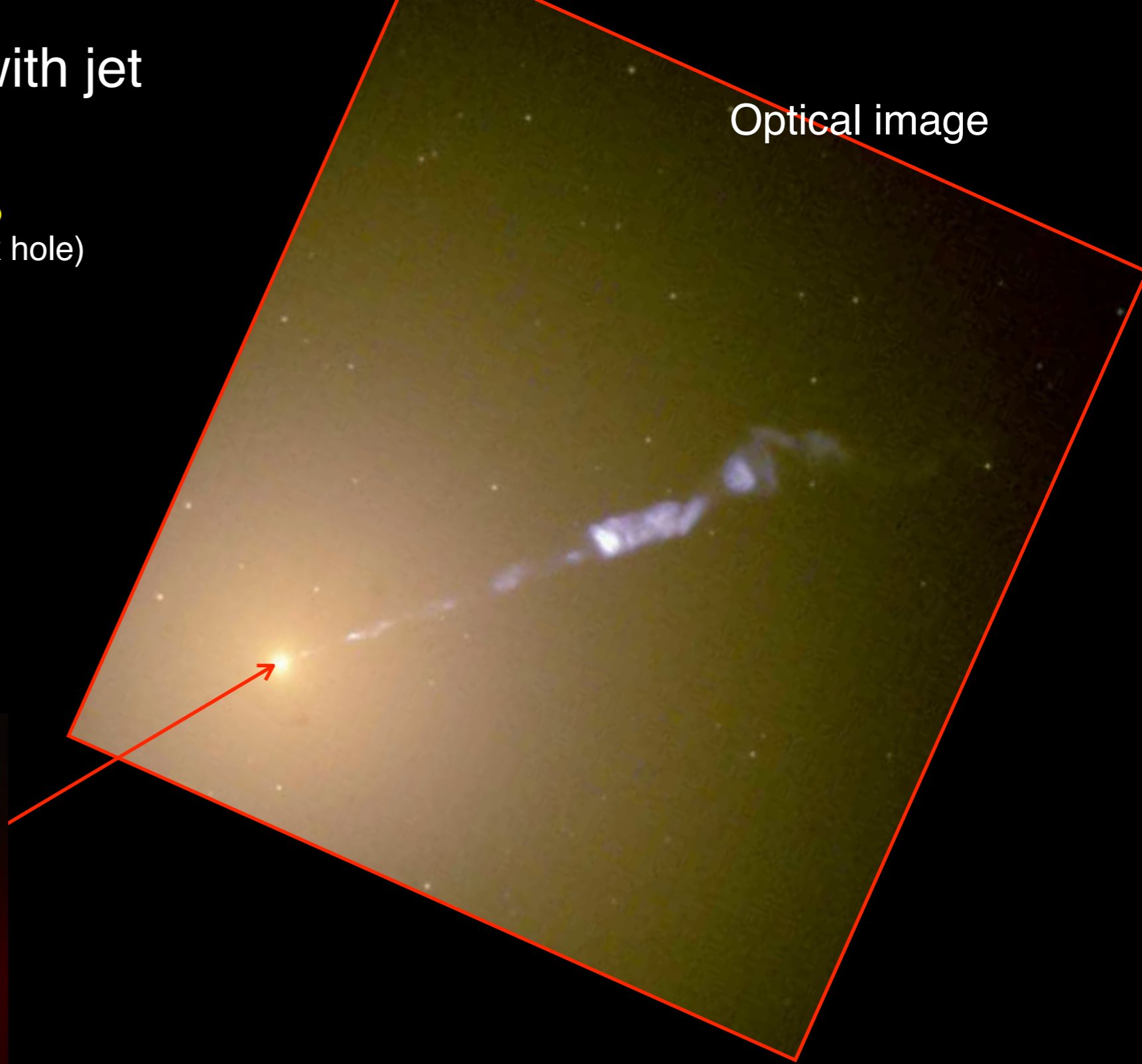
Total mass: $M_{tot} = 6 \times 10^{12} M_{\odot}$

Mass central black hole: $M_{BH} = 7 \times 10^9 M_{\odot}$

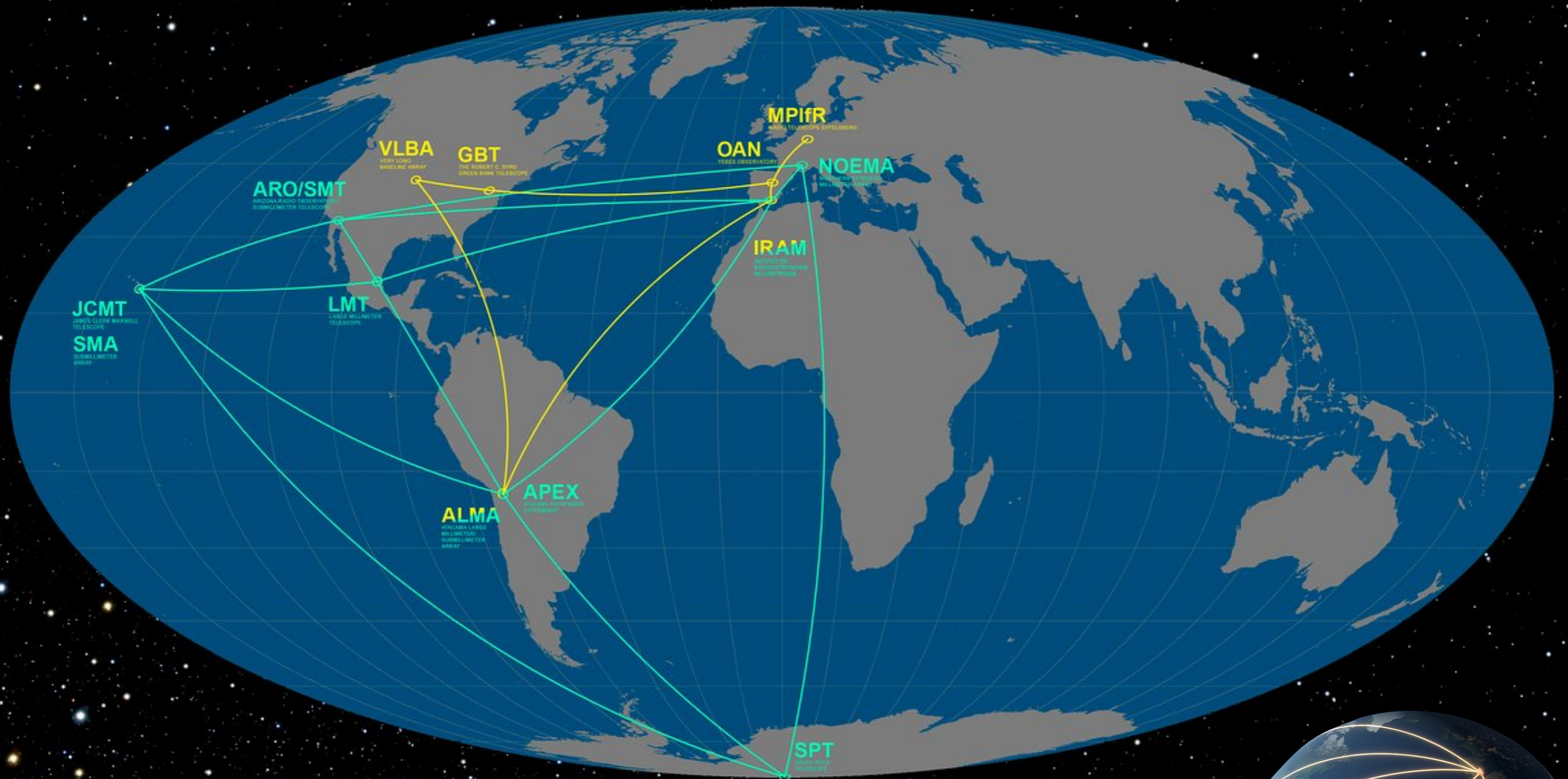
(one of the highest mass known for a black hole)

Optical image

Black hole at center



Event Horizon Telescope (EHT) to “observe” a black hole



EHT uses technique called *very-long-baseline interferometry* (VLBI)

8 radio observatories used on the entire Earth

Total data volume collected in 2017: 5 petabytes (350 terabytes/day)

Resolution achieved: 20 micro-arcseconds (20 μ as)

Wavelength observed: $\lambda = 1.3$ mm



Interferometry with relatively small radio antennas to get resolution of giant telescope

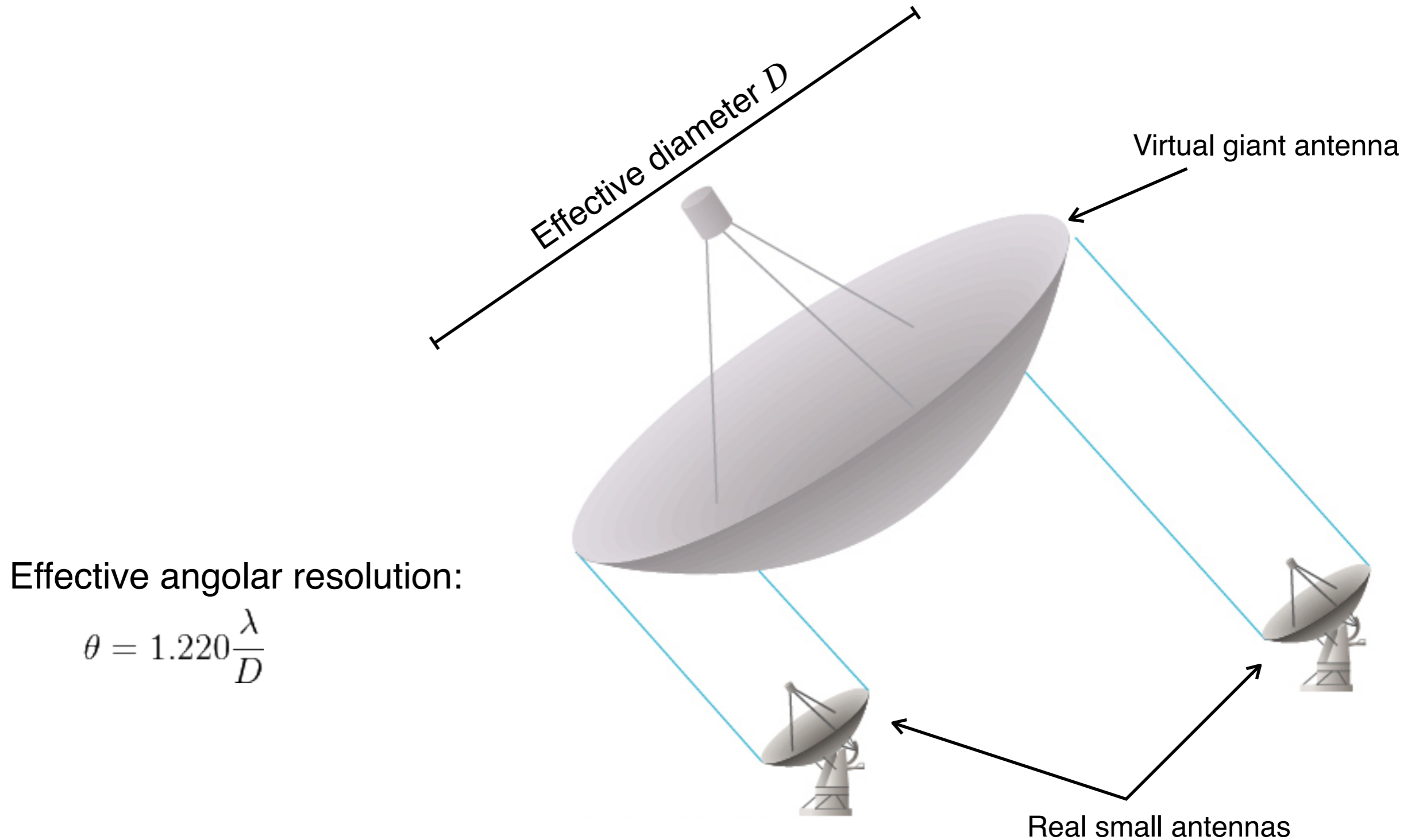


ALMA - Atacama Large Millimeter Array

66 12/7-meter diameter antennas

Movable antennas covering area from 150 metres to 16 kilometres

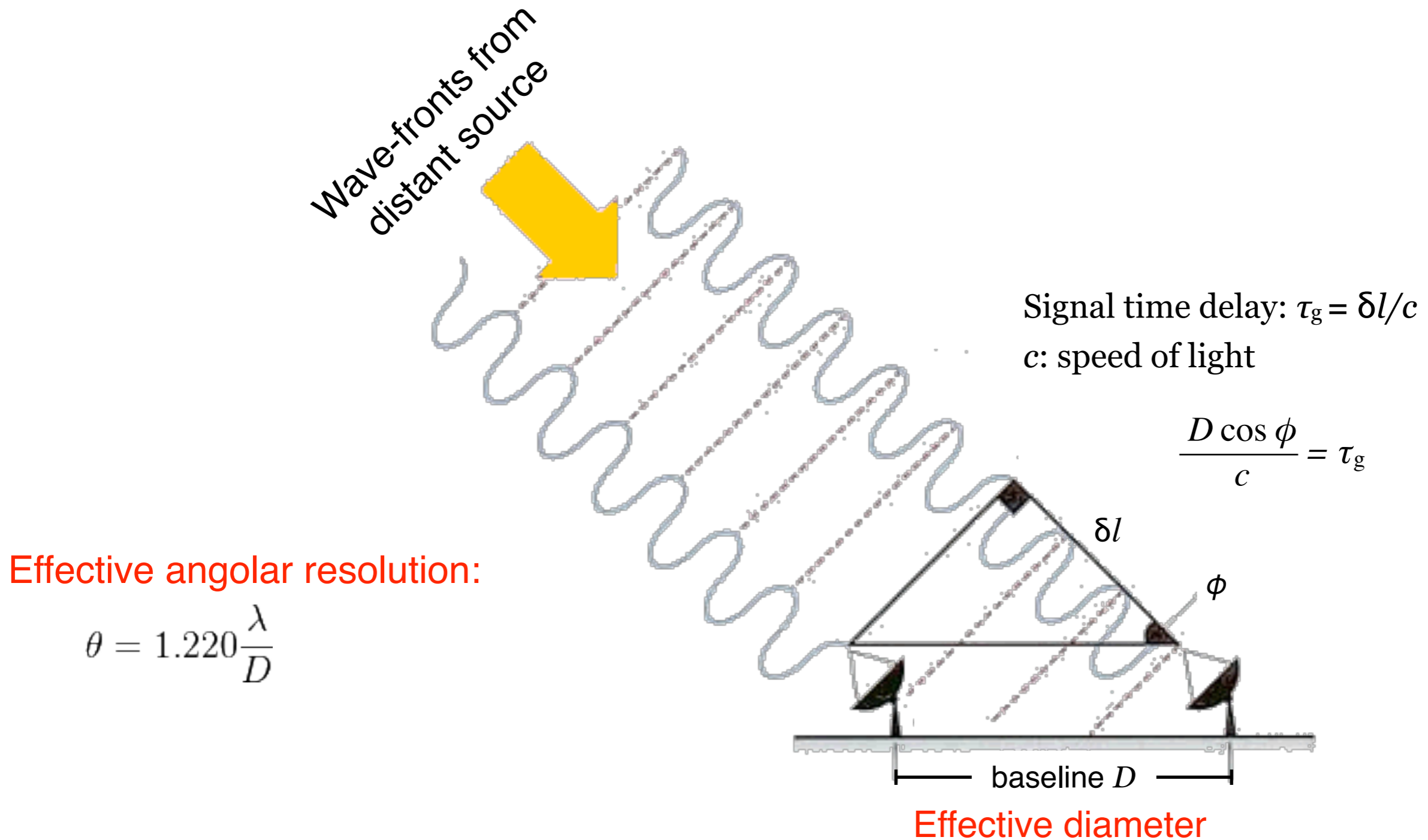
Interferometry with relatively small radio antennas to get resolution of giant telescope



Fringe pattern is created if separation between two antennas is $> 10 \times \lambda$

Interferometry for radio telescopes

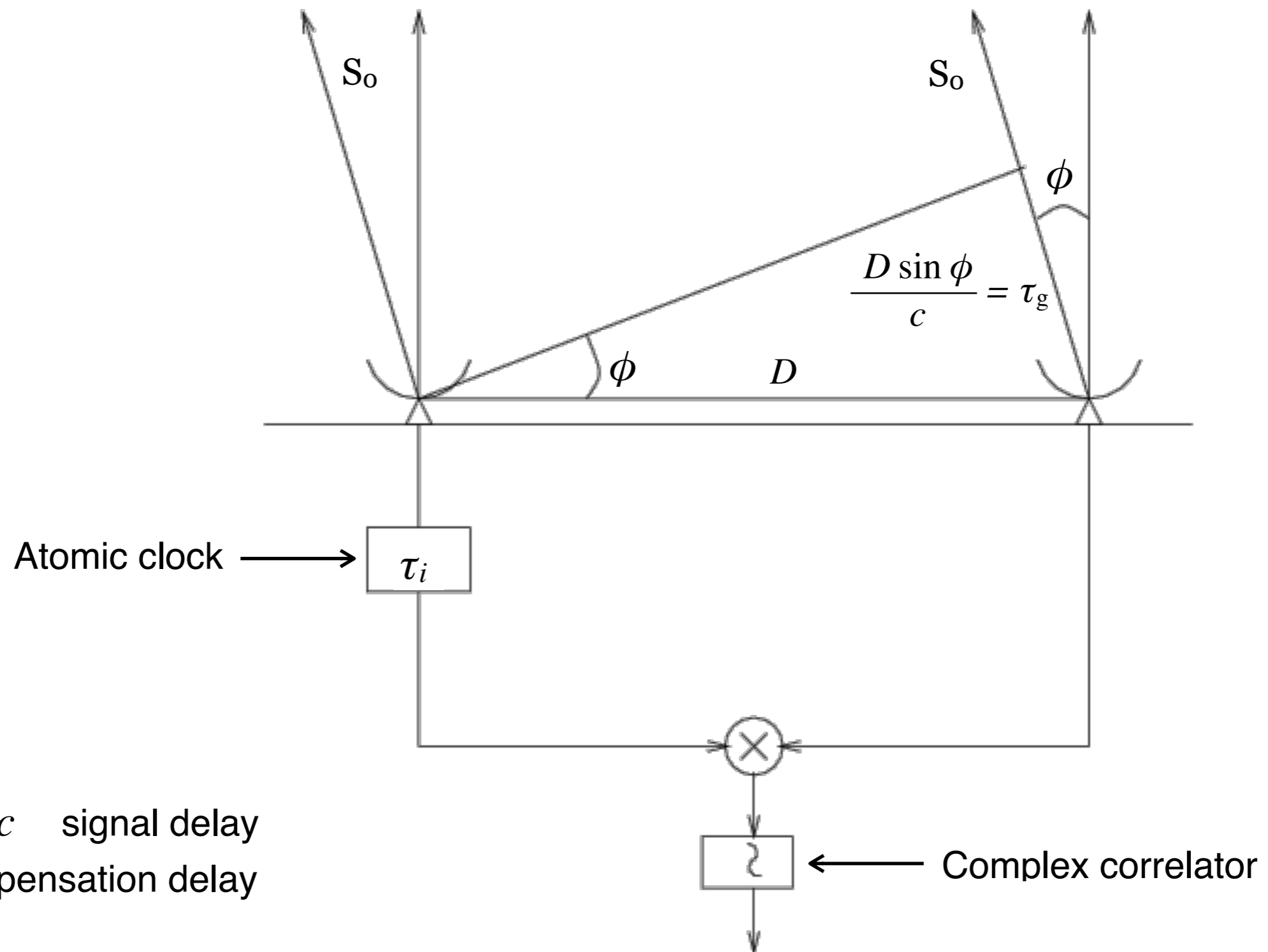
Interferometer: signal is combined from two or more telescopes to produce a sharper image to obtain **higher angular resolution**



Same source \implies signal perfectly in phase & same frequency

Interferometry for radio telescopes

Interferometer: signal is combined from two or more telescopes to produce a sharper image to obtain **higher angular resolution**

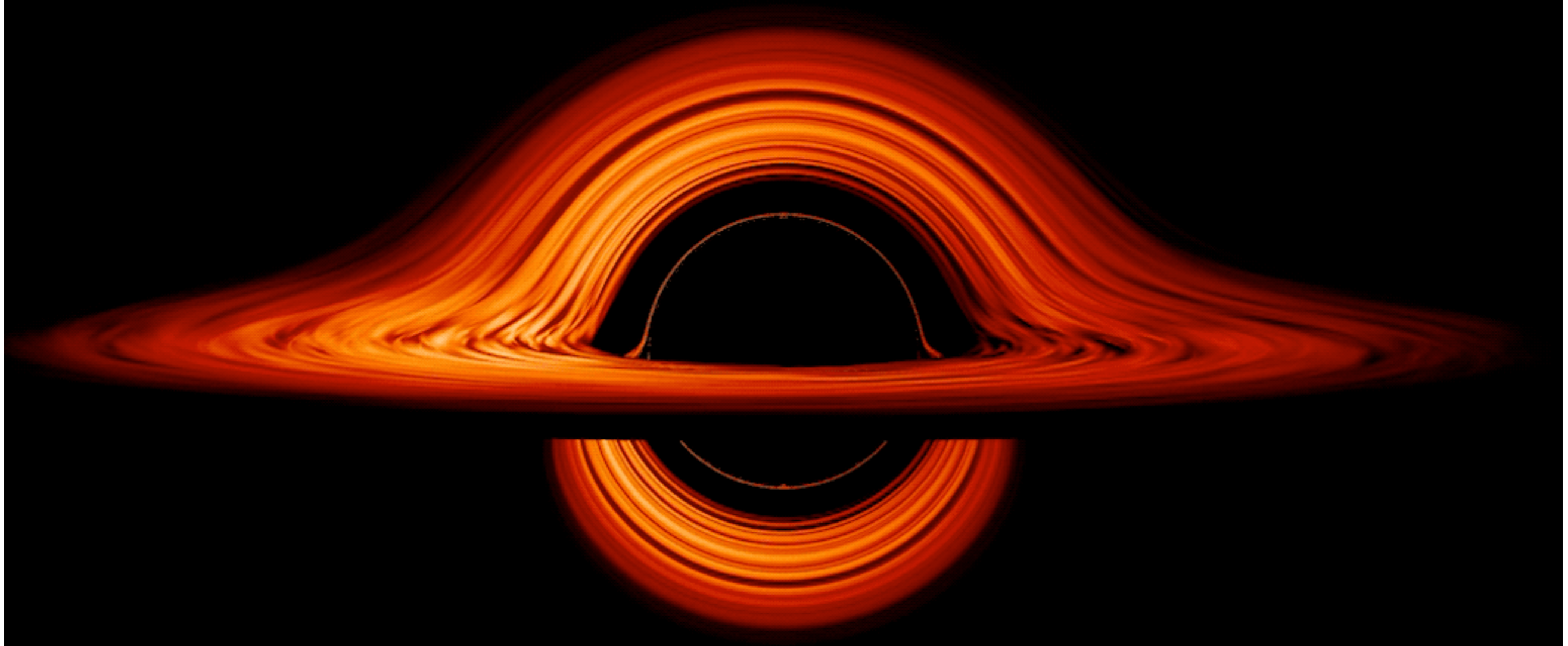


Most detailed image of a Black Hole ever obtained



Visualization of region around black hole as seen almost edge on

Computer simulation produced at NASA

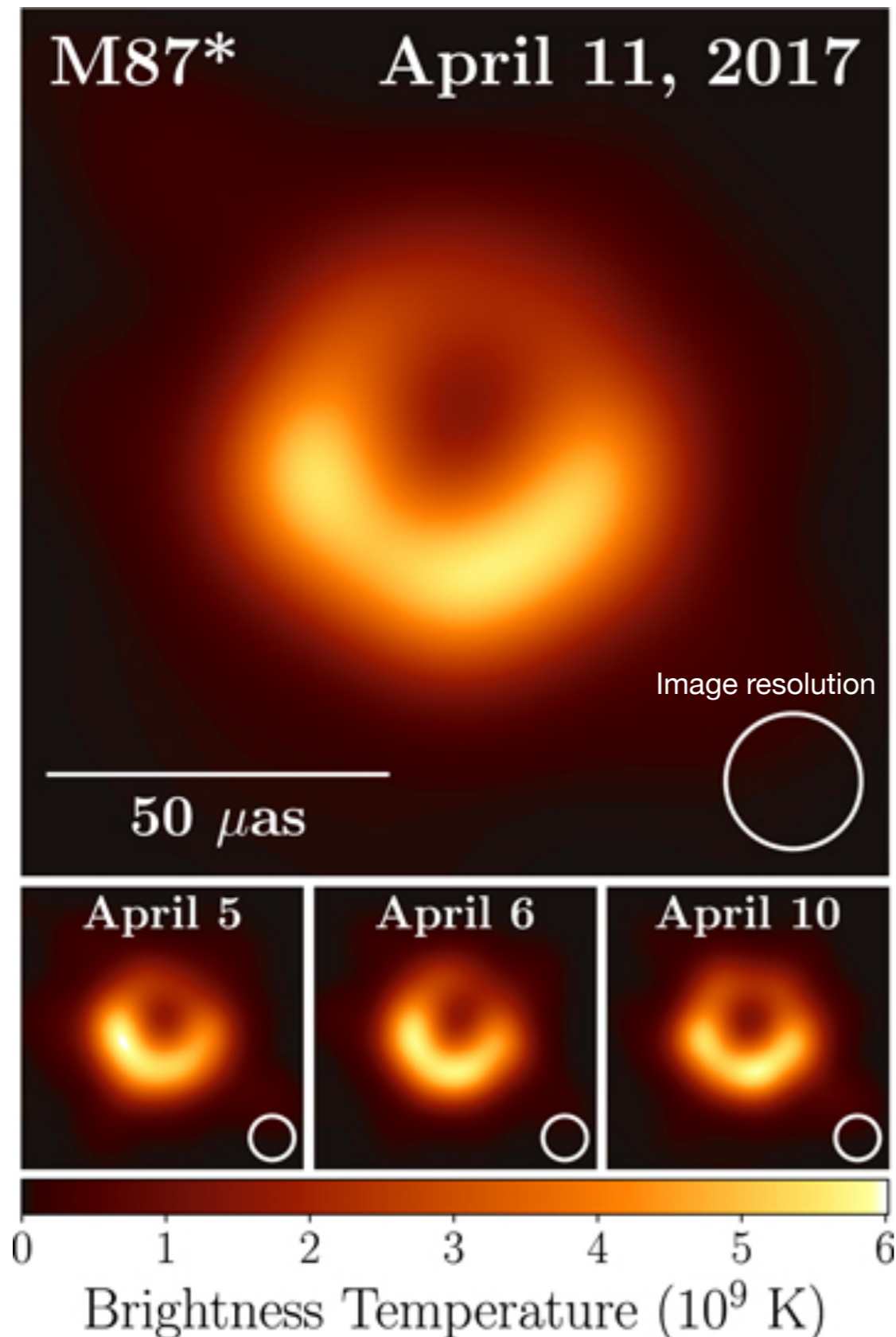


Relativistic effects:

- Disk emission distorted

Most detailed image of a Black Hole ever obtained

Direct studies of event horizon now possible via astronomical observations

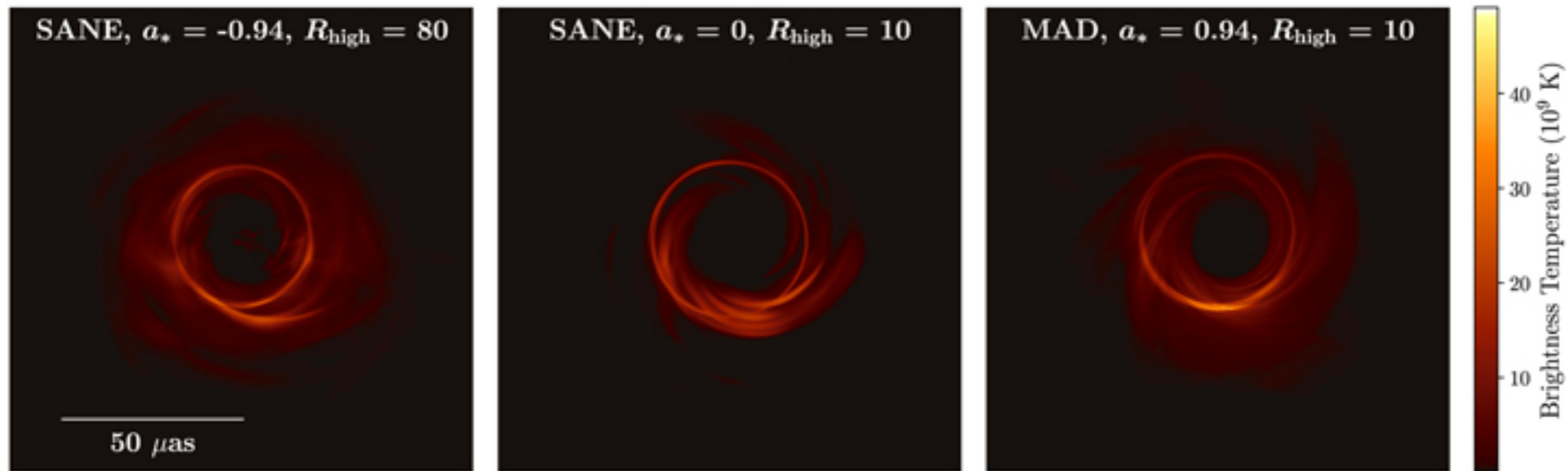


- Image resolution: 20 micro arc-seconds (μas)
- Equivalent to a **2 cents coin seen from Earth on the Moon**
- Completely dark region is where light cannot escape
- Luminous ring diameter: **$42 \pm 3 \mu\text{as}$** , brighter in the south
- Inclination angle of the orbiting disk with respect to Earth: 17°
- **Event horizon** around **2.5 times smaller than dark region**
- Size of event horizon is just under **40 billion km**
- Equivalent to the **orbit of Pluto around the Sun**
- Measured mass of black hole: $M_{BH} = (6.5 \pm 0.7) \times 10^9 M_\odot$

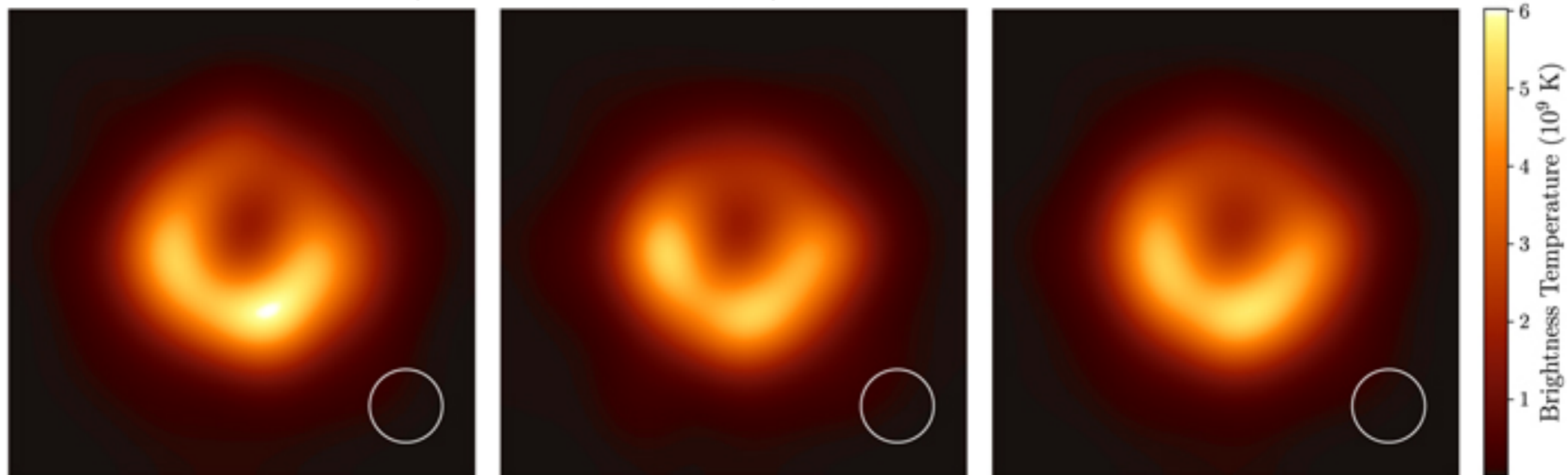
Theoretical expectations & computer simulations

40-50 thousand simulations, a few reproducing observations

Theoretical models of accretion flows



Computer simulation of expected observations



Future: **Space** Very Large Baseline Interferometry

Bibliography: 6 papers published in *Astrophysical Journal Letters*

- Paper I: [The Shadow of the Supermassive Black Hole](#)
- Paper II: [Array and Instrumentation](#)
- Paper III: [Data processing and Calibration](#)
- Paper IV: [Imaging the Central Supermassive Black Hole](#)
- Paper V: [Physical Origin of the Asymmetric Ring](#)
- Paper VI: [The Shadow and Mass of the Central Black Hole](#)